

(4.8) The interior of the earth is assumed to be built up of a solid core (radius 1371 km) followed by a molten core (radius 3471 km) and a molten mantle (radius 6354 km) covered by a 17 km thick crust. One assumes that 2% by weight of the molten mantle and crust is potassium; the mantle density is assumed to be 6000 kg m⁻³ and that of the crust 3300 kg m⁻³. What energy outflow will the radioactive decay of this element cause at the earth's surface? The decay scheme of ⁴⁰K is given in eqn. (12.16). For the EC branch Q = 1.505 MeV, for the beta-branch 1.314 MeV. Compare this energy output to the solar energy influx to the earth of 3.2x10²⁴ J y⁻¹.

$$Sunflux := 3.2 \cdot 10^{24} \cdot \text{joule} \cdot \text{yr}^{-1}$$

$$N_A := 6.022137 \cdot 10^{23} \cdot \text{mole}^{-1}$$

$$r_y := 6354 \cdot 10^3 \cdot \text{m}$$

$$r_i := 3471 \cdot 10^3 \cdot \text{m}$$

$$r_{cr} := r_y + 17000 \cdot \text{m}$$

$$d := 5000 \cdot \text{kg} \cdot \text{m}^{-3}$$

$$d_{cr} := 3300 \cdot \text{kg} \cdot \text{m}^{-3}$$

$$V_{mantle} := \frac{4}{3} \cdot \pi \cdot (r_y^3 - r_i^3)$$

$$M_{mantle} := V_{mantle} \cdot d$$

$$M_{mantle} = 4.497 \cdot 10^{24} \cdot \text{kg}$$

$$V_{crust} := \frac{4}{3} \cdot \pi \cdot (r_{cr}^3 - r_y^3)$$

$$M_{crust} := V_{crust} \cdot d_{cr}$$

$$M_{crust} = 2.854 \cdot 10^{22} \cdot \text{kg}$$

$$M_K := (M_{mantle} + M_{crust}) \cdot \frac{2}{100}$$

$$M_{K40} := M_K \cdot \frac{0.0117}{100}$$

$$M_{K40} = 1.059 \cdot 10^{19} \cdot \text{kg}$$

$$M_{wK40} := 40 \cdot \text{gm} \cdot \text{mole}^{-1}$$

$$N_{K40} := \frac{M_{K40}}{M_{wK40}} \cdot N_A$$

$$t_{half} := 1.28 \cdot 10^9 \cdot \text{yr}$$

$$\lambda := \frac{\ln(2)}{t_{half}}$$

$$\lambda = 1.716 \cdot 10^{-17} \cdot \text{sec}^{-1}$$

$$R_{K40} := \lambda \cdot N_{K40}$$

$$R_{K40} = 2.736 \cdot 10^{27} \cdot \text{sec}^{-1}$$

If we include the energy of escaping neutrinos, i.e. use total decay energies, the energy production rate is:

$$E_{decay} := (0.1067 \cdot 1.505 + 0.8933 \cdot 1.312) \cdot 10^6 \cdot 1.6021773 \cdot 10^{-19} \cdot \text{joule}$$

$$E_1 := R_{K40} \cdot E_{decay} \quad E_1 = 1.843 \cdot 10^{22} \cdot \frac{\text{joule}}{\text{yr}}$$

$$Y := 4 \cdot \pi \cdot r_{cr}^2 \quad Y = 5.101 \cdot 10^{14} \cdot \text{m}^2 \quad \frac{E_1}{Y} = 1.145 \cdot \text{watt} \cdot \text{m}^{-2} \quad \frac{E_1}{Sunflux} = 0.5760 \cdot \%$$

If we consider energy loss by escaping neutrinos, i.e. use average β -energies from Table of Isotopes, we obtain:

$$Q_{eff} := (.1561 + .455 + 0.00062) \cdot 10^6 \cdot 1.6021773 \cdot 10^{-19} \cdot \text{joule}$$

$$E_2 := R_{K40} \cdot Q_{eff} \quad E_2 = 8.462 \cdot 10^{21} \cdot \frac{\text{joule}}{\text{yr}} \quad \frac{E_2}{Y} = 0.526 \cdot \text{m}^{-2} \cdot \text{watt}$$

Solar influx 1.94 cal/cm² min

$$k_{solar} := 1.94 \cdot \text{cal} \cdot \text{cm}^{-2} \cdot \text{min}^{-1}$$

$$k_{solar} = 1.354 \cdot 10^3 \cdot \text{m}^{-2} \cdot \text{watt}$$

$$A := \pi \cdot r_{cr}^2 \cdot k_{solar}$$

$$A = 1.726 \cdot 10^{17} \cdot \text{sec}^{-1} \cdot \text{joule}$$

$$P_{heat} := A \cdot 0.58$$

$$r_{cr} = 6.371 \cdot 10^6 \cdot \text{m}$$

$$A \cdot 0.58 = 1.001 \cdot 10^{17} \cdot \text{sec}^{-1} \cdot \text{joule}$$

(J/y becomes heat)

$$P_{heat} = 3.16 \cdot 10^{24} \cdot \frac{\text{joule}}{\text{yr}}$$