

(5.8) The interior of the earth is assumed to be built up of a solid core (radius 1371 km) followed by a molten core (radius 3471 km) and a molten mantle (radius 6354 km) covered by a 17 km thick crust. One assumes that 2% by weight of the molten mantle and crust is potassium; the mantle density is assumed to be 6000 kg m⁻³ and that of the crust 3300 kg m⁻³. What energy outflow will the radioactive decay of this element cause at the earth's surface? The decay scheme of ⁴⁰K is given in eqn. (13.7). For the EC branch Q = 1.505 MeV, for the beta-branch 1.314 MeV. Compare this energy output to the solar energy influx to the earth of 3.2x10²⁴ J y⁻¹.

$$\begin{aligned}
 \text{Sunflux} &:= 3.2 \cdot 10^{24} \cdot \text{joule} \cdot \text{yr}^{-1} & N_A &:= 6.022137 \cdot 10^{23} \cdot \text{mole}^{-1} \\
 r_y &:= 6354 \cdot 10^3 \cdot \text{m} & r_i &:= 3471 \cdot 10^3 \cdot \text{m} & r_{cr} &:= r_y + 17000 \cdot \text{m} \\
 d &:= 5000 \cdot \text{kg} \cdot \text{m}^{-3} & d_{cr} &:= 3300 \cdot \text{kg} \cdot \text{m}^{-3} \\
 V_{\text{mantle}} &:= \frac{4}{3} \cdot \pi \cdot (r_y^3 - r_i^3) & M_{\text{mantle}} &:= V_{\text{mantle}} \cdot d & M_{\text{mantle}} &= 4.497 \cdot 10^{24} \cdot \text{kg} \\
 V_{\text{crust}} &:= \frac{4}{3} \cdot \pi \cdot (r_{cr}^3 - r_y^3) & M_{\text{crust}} &:= V_{\text{crust}} \cdot d_{cr} & M_{\text{crust}} &= 2.854 \cdot 10^{22} \cdot \text{kg} \\
 M_K &:= (M_{\text{mantle}} + M_{\text{crust}}) \cdot \frac{2}{100} & M_{K40} &:= M_K \cdot \frac{0.0117}{100} & M_{K40} &= 1.059 \cdot 10^{19} \cdot \text{kg} \\
 M_{wK40} &:= 40 \cdot \text{gm} \cdot \text{mole}^{-1} \\
 N_{K40} &:= \frac{M_{K40}}{M_{wK40}} \cdot N_A & t_{\text{half}} &:= 1.28 \cdot 10^9 \cdot \text{yr} & \lambda &:= \frac{\ln(2)}{t_{\text{half}}} & \lambda &= 1.716 \cdot 10^{-17} \cdot \text{sec}^{-1} \\
 R_{K40} &:= \lambda \cdot N_{K40} & R_{K40} &= 2.736 \cdot 10^{27} \cdot \text{sec}^{-1}
 \end{aligned}$$

If we include the energy of escaping neutrinos, i.e. use total decay energies, the energy production rate is:

$$\begin{aligned}
 E_{\text{decay}} &:= (0.1067 \cdot 1.505 + 0.8933 \cdot 1.312) \cdot 10^6 \cdot 1.6021773 \cdot 10^{-19} \cdot \text{joule} \\
 E_1 &:= R_{K40} \cdot E_{\text{decay}} & E_1 &= 1.843 \cdot 10^{22} \cdot \frac{\text{joule}}{\text{yr}} \\
 Y &:= 4 \cdot \pi \cdot r_{cr}^2 & Y &= 5.101 \cdot 10^{14} \cdot \text{m}^2 & \frac{E_1}{Y} &= 1.145 \cdot \text{watt} \cdot \text{m}^{-2} & \frac{E_1}{\text{Sunflux}} &= 0.5760 \cdot \%
 \end{aligned}$$

If we consider energy loss by escaping neutrinos, i.e. use average β -energies from Table of Isotopes, we obtain:

$$\begin{aligned}
 Q_{\text{eff}} &:= (.1561 + .455 + 0.00062) \cdot 10^6 \cdot 1.6021773 \cdot 10^{-19} \cdot \text{joule} \\
 E_2 &:= R_{K40} \cdot Q_{\text{eff}} & E_2 &= 8.462 \cdot 10^{21} \cdot \frac{\text{joule}}{\text{yr}} & \frac{E_2}{Y} &= 0.526 \cdot \text{m}^{-2} \cdot \text{watt}
 \end{aligned}$$

Solar influx 1.94 cal/cm² min

$$\begin{aligned}
 k_{\text{solar}} &:= 1.94 \cdot \text{cal} \cdot \text{cm}^{-2} \cdot \text{min}^{-1} & k_{\text{solar}} &= 1.354 \cdot 10^3 \cdot \text{m}^{-2} \cdot \text{watt} \\
 A &:= \pi \cdot r_{cr}^2 \cdot k_{\text{solar}} & A &= 1.726 \cdot 10^{17} \cdot \text{sec}^{-1} \cdot \text{joule} & P_{\text{heat}} &:= A \cdot 0.58 \\
 r_{cr} &= 6.371 \cdot 10^6 \cdot \text{m} & A \cdot 0.58 &= 1.001 \cdot 10^{17} \cdot \text{sec}^{-1} \cdot \text{joule} & & (\text{J/y becomes heat}) \\
 P_{\text{heat}} &= 3.16 \cdot 10^{24} \cdot \frac{\text{joule}}{\text{yr}}
 \end{aligned}$$